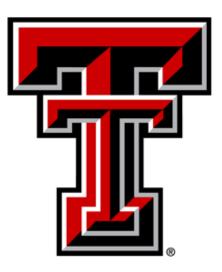
WTLMA, LF, and GLM comparisons with updated processing

GLM Science Meeting

23 September 2021

Eric C. Bruning Texas Tech University, Lubbock, TX

Acknowledgments: Thanks to Jason Jordan, Stephanie Weiss, and David PeQueen and many colleagues for their contributions to last year's version of this talk. This work supported by NASA award 80NSSC19K1576, and NOAA JTTI award NA19NES4320002 via CISESS-Maryland. Earth Networks, Inc. and Vaisala Inc. provided updated datasets that made this processing comparison possible.



Purpose Assess updated ground network datasets

Last year: showed significant differences in cloud and ground stroke detection efficiency and classification on 9 Oct 2019 in Lubbock, TX.

Advocated for: a long term effort to crosscheck and validate detection efficiency and classification performance, since innovation means ground datasets continue to change.

Both classification/location algorithms were since updated: are the changes meaningful?

TEXAS TECH UNIV

Datasets:

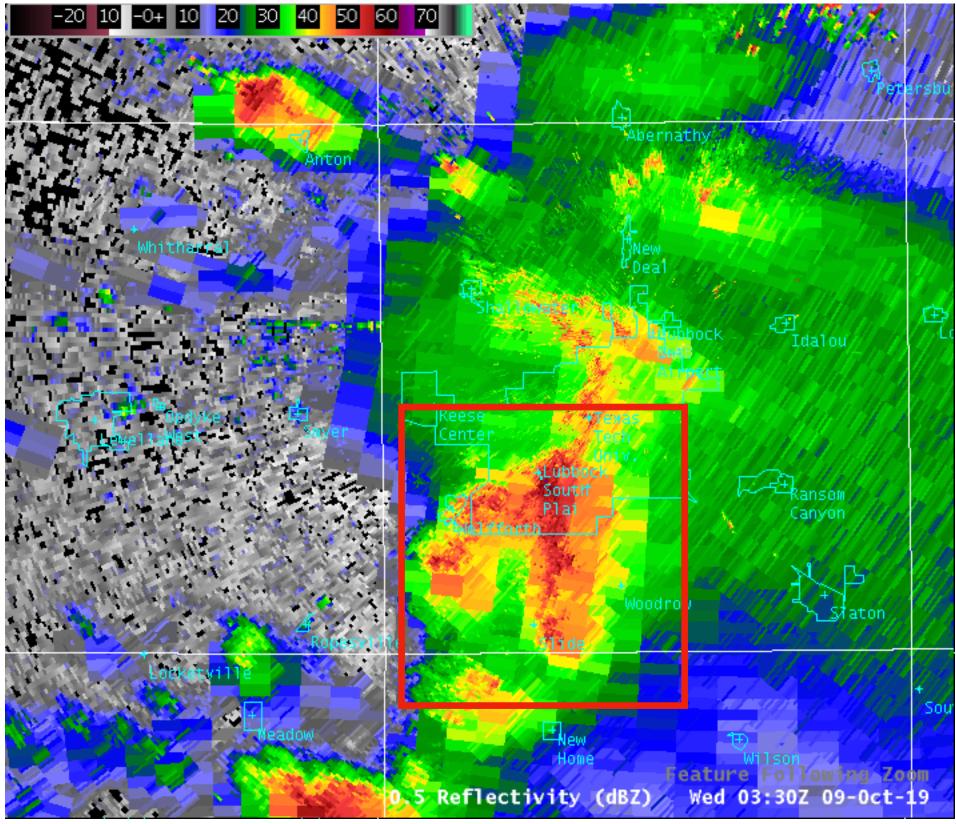
- West Texas Lightning Mapping Array (VHF band)
- GOES-16, GOES-17 GLM (optical)
- ENI and Vaisala stroke-detecting networks (~LF band) — reprocessed datasets (thanks!)



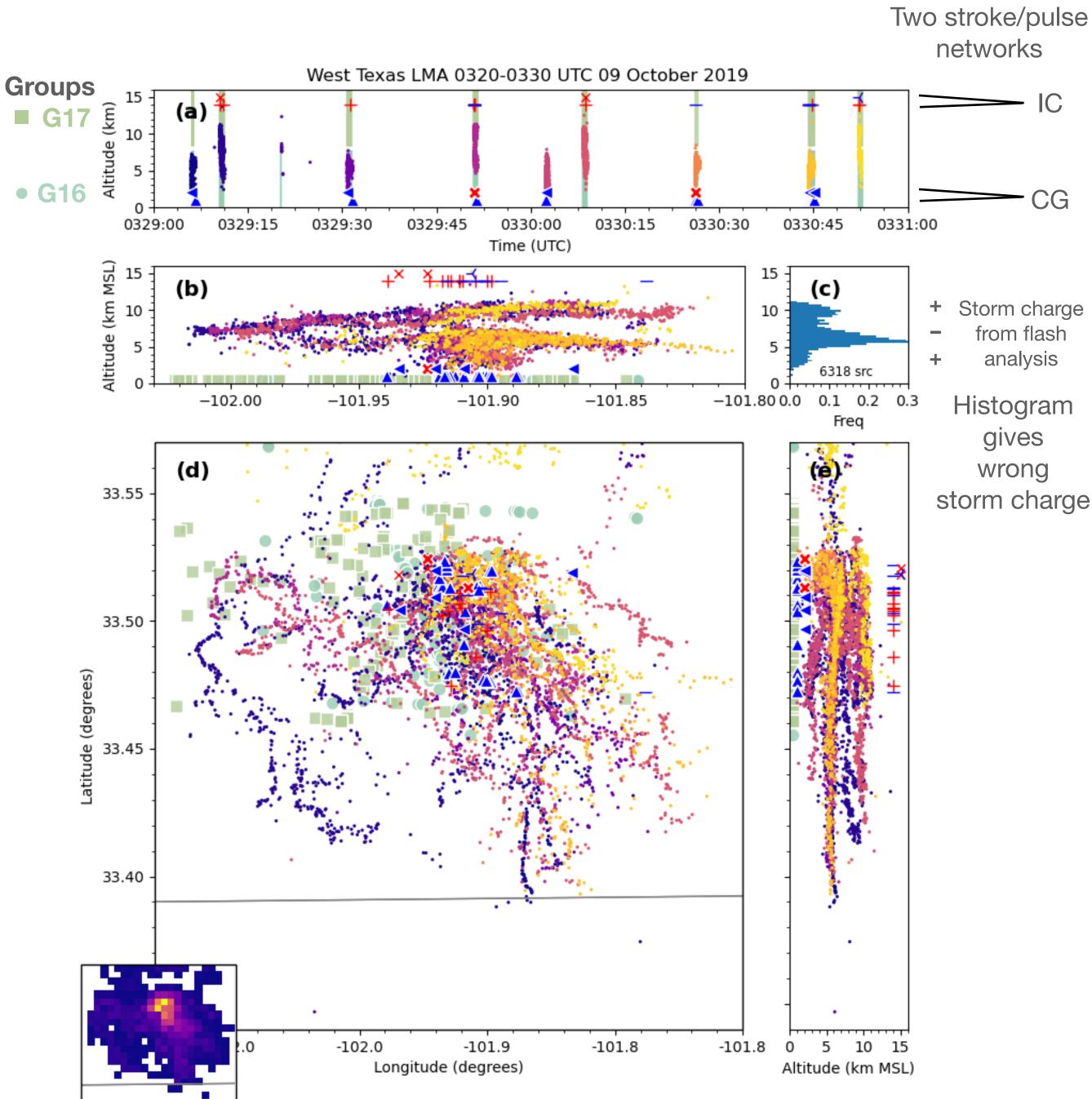
Classification: mixed messages

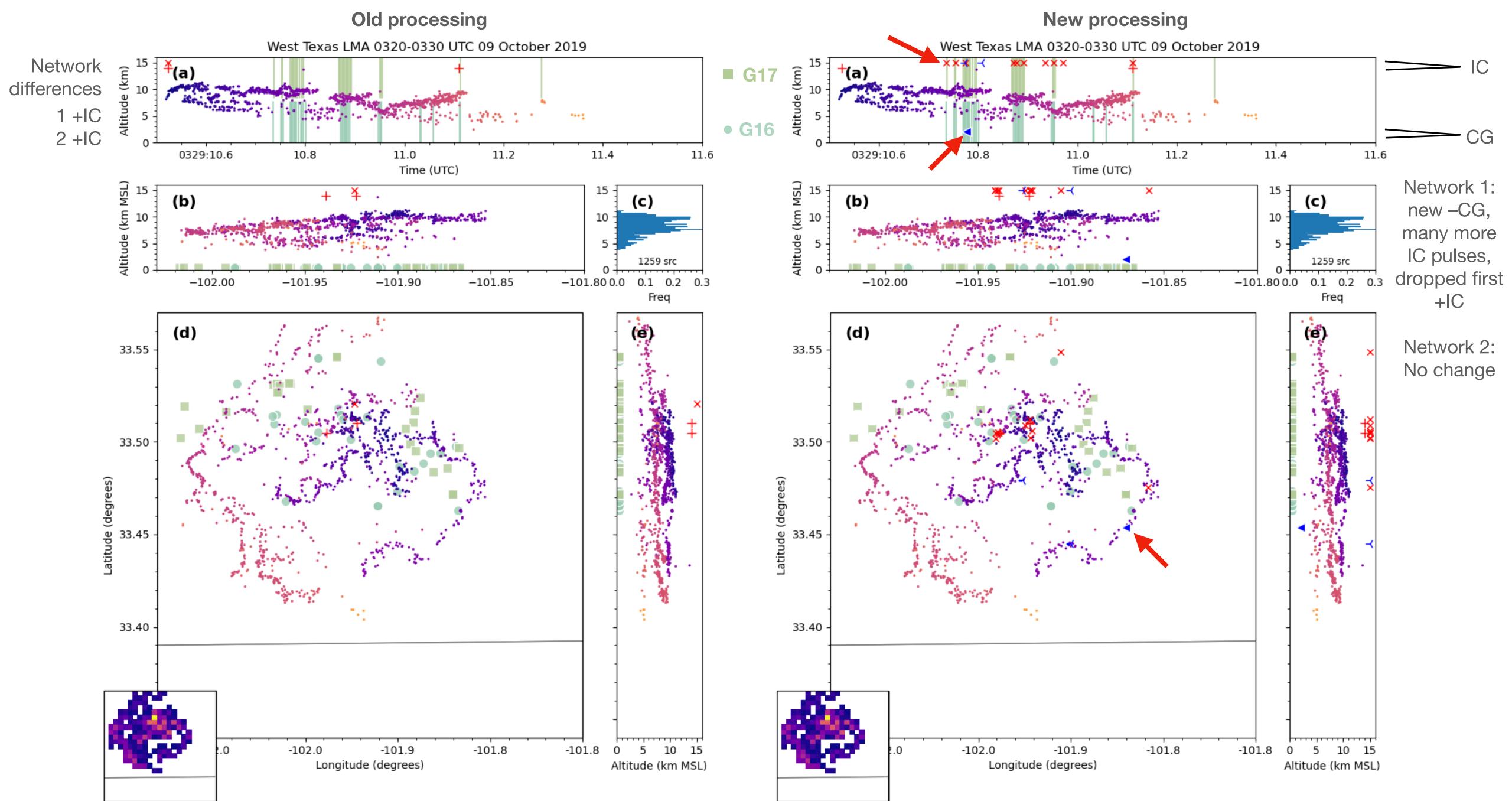
Last year: 9 flashes, two minutes, quite a bit of disagreement

• Case on 8-9 October observed by coauthor Jordan: saw two lightning flashes out window and looked at two radiofrequency stroke detection networks for their solutions.



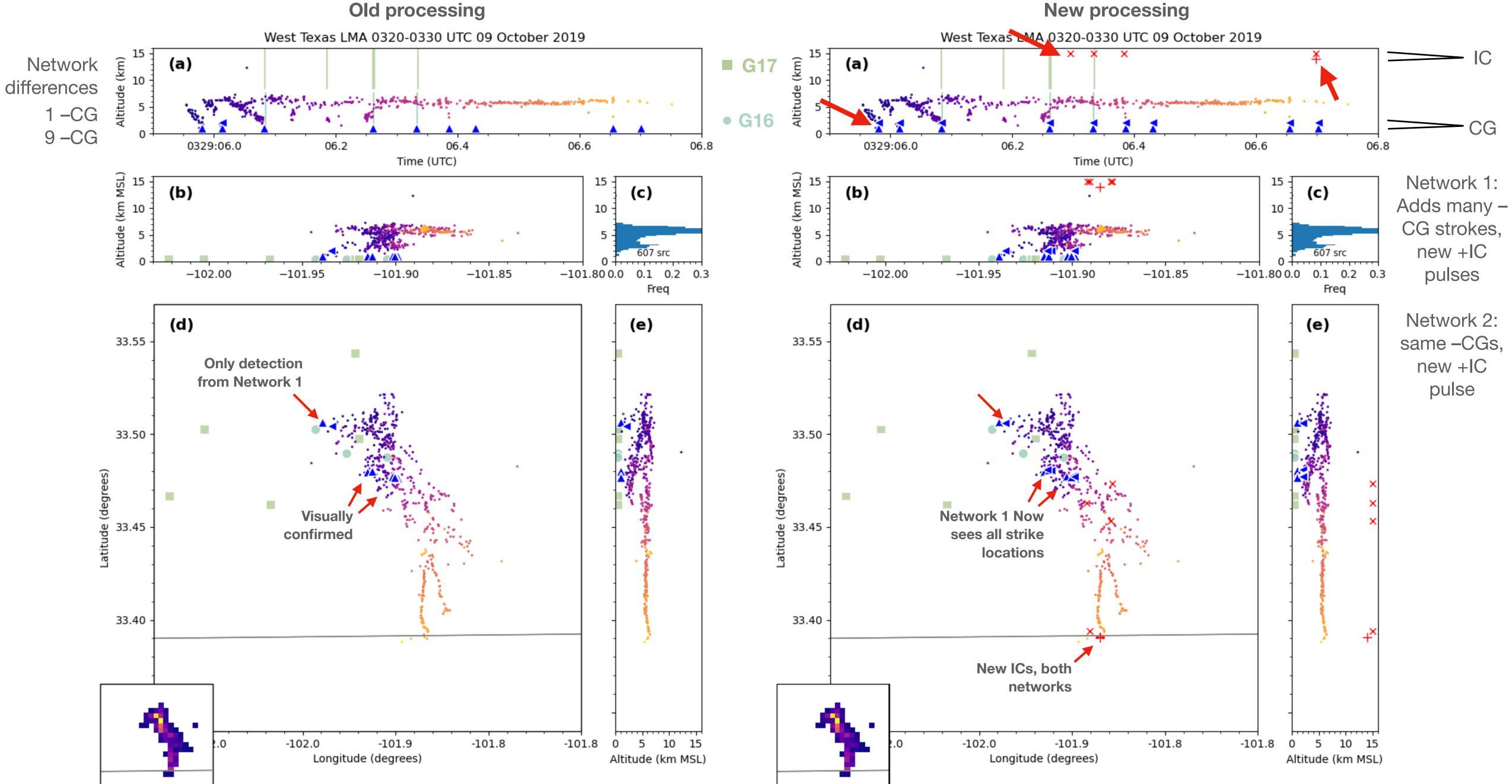
Radar reflectivity

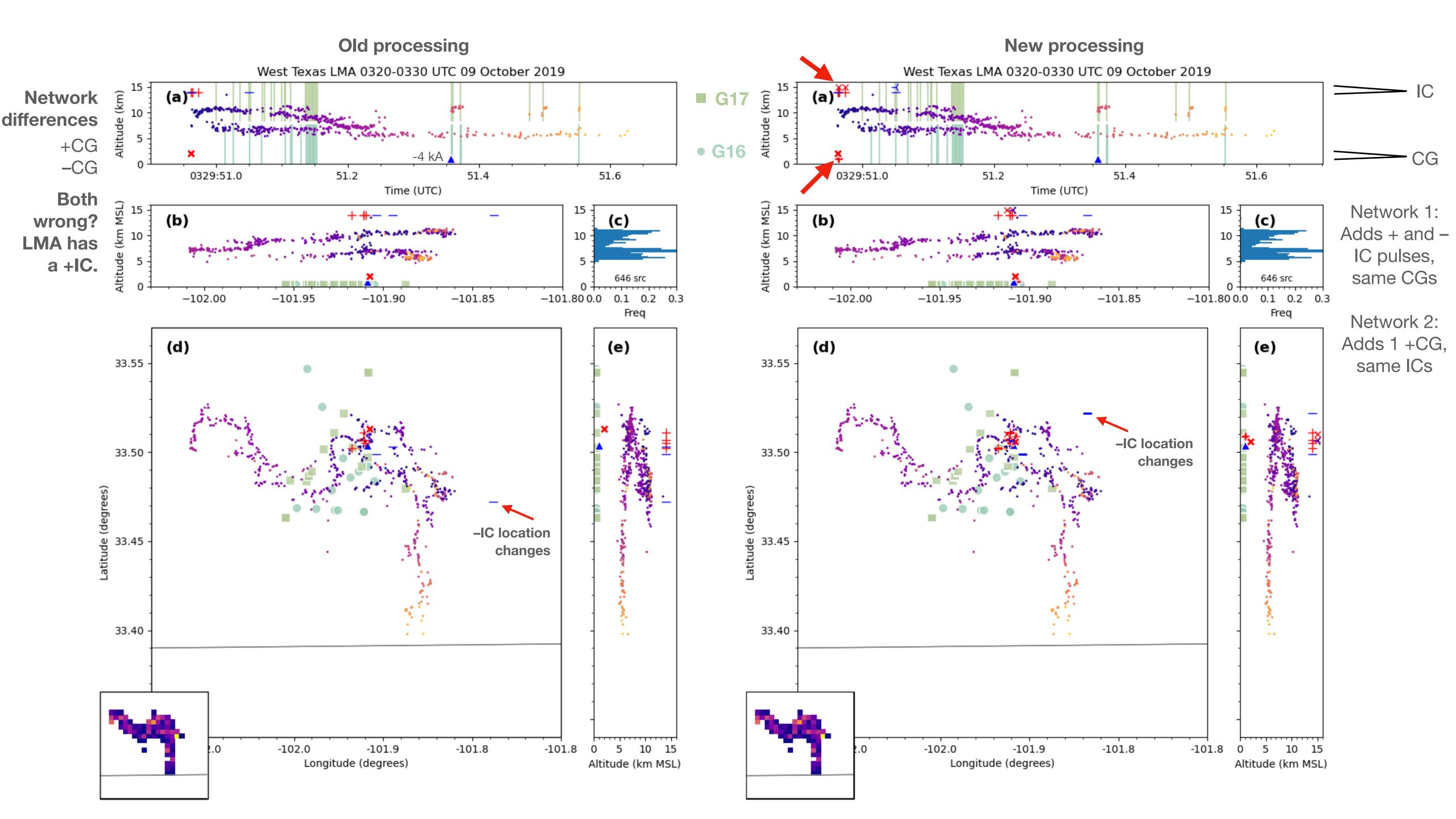


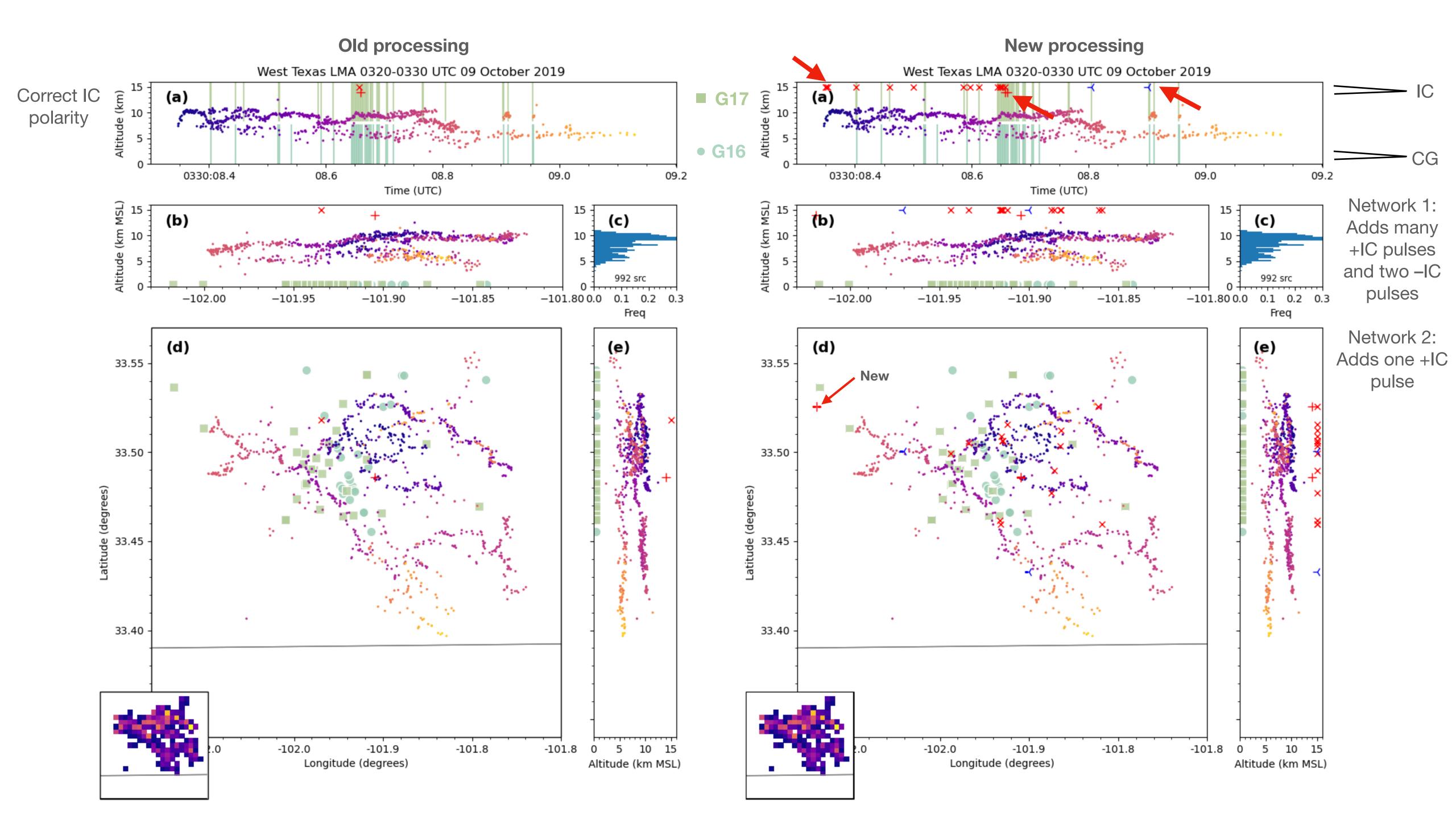


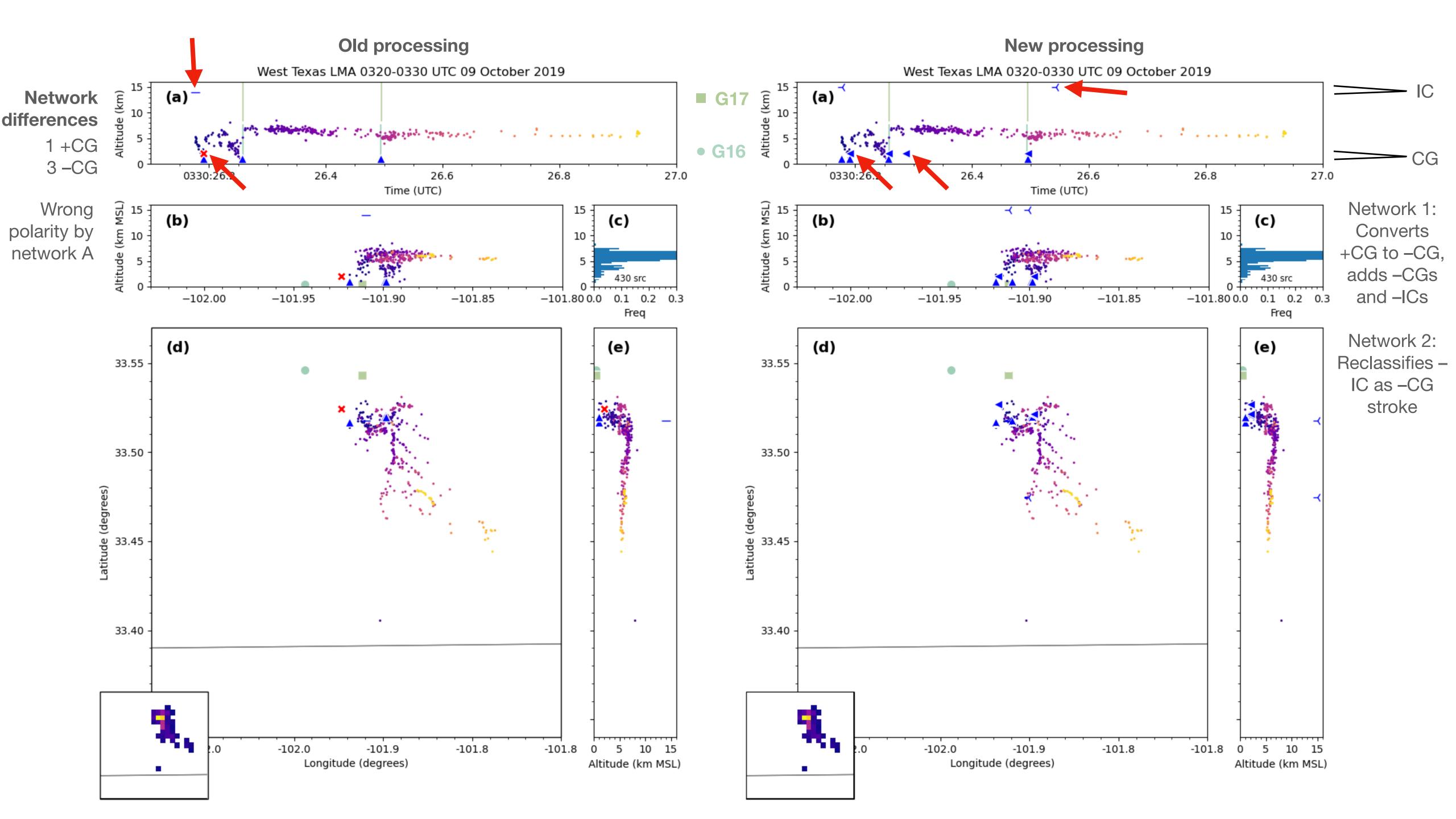
Good overall IC performance by all sensors

CGs close to Jason Jordan's house









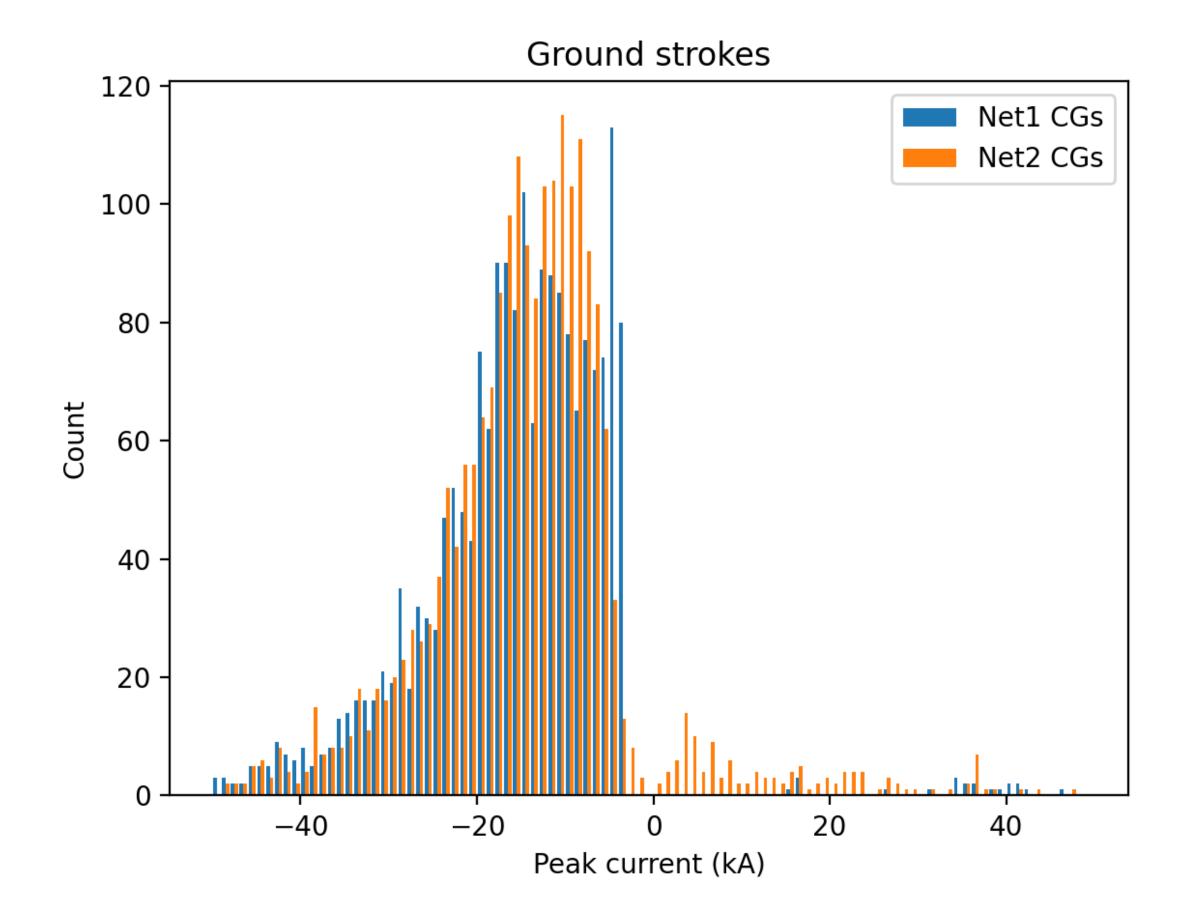
WTLMA, LF, and GLM compa Summary

- Limited sample; no field change data for cross-validation.
- Processing changes have brought both datasets into closer agreement. Network 1 changed a lot, small differences in network 2.
- Differences in classification and errors in location are still present. QC data (# sensors, error ellipses) increasingly important for flexible filtering of low amplitude events.

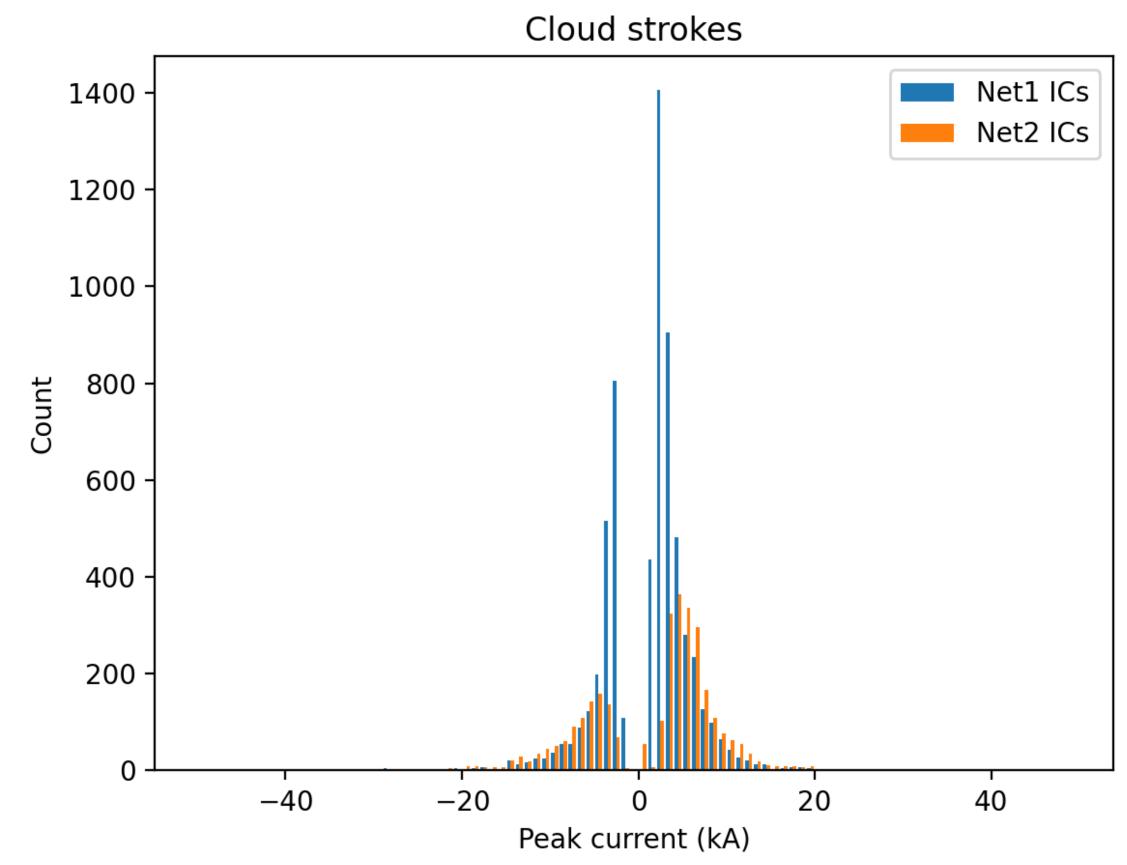
TEXAS TECH UNIV WTLMA, LF, and GLM comparisons with updated processing

- Mixed IC polarities and/or IC polarity that disagrees with CG polarity is not uncommon. Utilization of IC polarity requires a deeper background in lightning than we can expect of meteorologists?
- I still advocate for a commitment to a geographically diverse set of field change validation data to complement LMAs — ideally public and open.

Peak current distribution With new processing



0300-0400 UTC, 33 to 35° latitude, -103 to -101° longitude



Preparatory work for MTG LI accumulated products

- Added code to glmtools to calculate illuminated flash fraction (IFF), a baseline MTG LI accumulated product.
 - Sum of grid is total flash rate
 - Pixels are illuminated in proportion to the fraction of time they were illuminated during the flash
- Created categorized bibliography of all 233 papers mentioning GLM from 2018 – June 2021 in AMS and AGU journals.
- Created animations in overlap region of GLM and MTG LI (next)

1 min

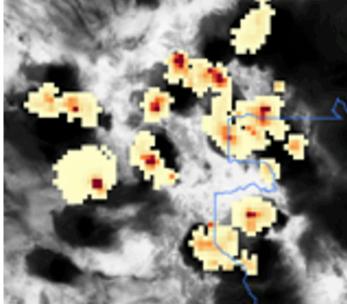
5 min

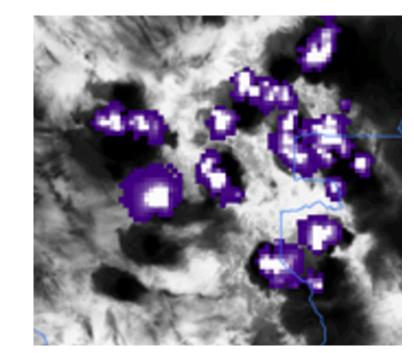


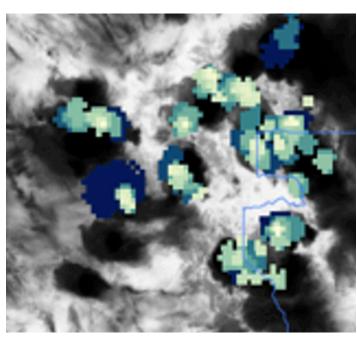


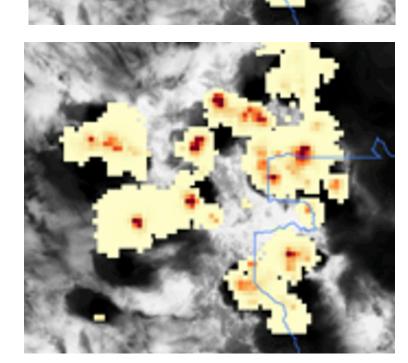


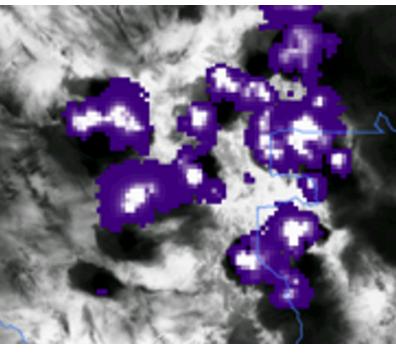


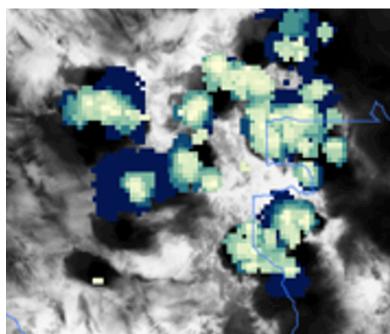












TOE

MFA



5 min window animations

GOES ABI Ch 13 (10 μm thermal IR) and GLM

- Overlays were constructed with SatPy, over each subregion and for all four variables.
- 00 UTC, 1 September to 08 UTC, 4 September 2018.
- Individual images and loops available at the website, <u>http://pogo.tosm.ttu.edu/data/</u> <u>MTG/GLM_MTG_proxy_2018/</u> region_images/

